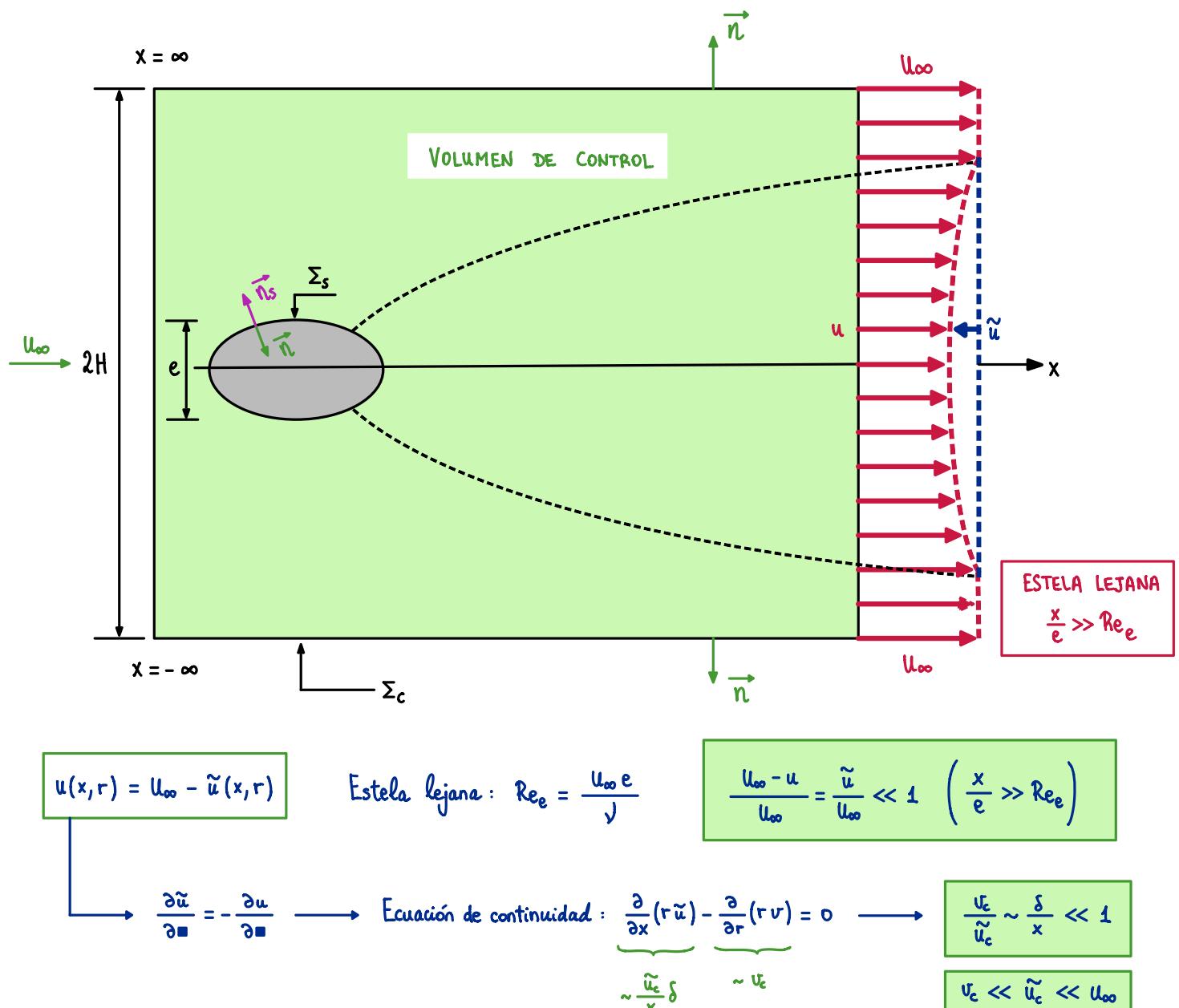


# Estela Axilísmetrica Láminar



Ecuación de continuidad aplicada al volumen de control:

$$\int_{\Sigma_c} \vec{v} \cdot \vec{n} dA = 0 \longrightarrow -u_\infty \frac{\pi(2H)^2}{4} + 2 \int_{-\infty}^x v \pi H dx + 2 \int_0^H u \pi r dr = 0$$



ECdM<sub>x</sub> aplicada al volumen de control:

$$\int_{\Sigma_c} \rho \vec{v} (\vec{v} \cdot \vec{n}) dA \vec{i} = - \int_{\Sigma_s} (\rho - \rho_\infty) \vec{n} dA \vec{i} + \int_{\Sigma_c} \vec{\tau}_u \cdot \vec{n} dA \vec{i}$$

$\underbrace{-D}_{-D}$

$$-\bar{U}_\infty \frac{\pi(2H)^2}{4} + 2\bar{U}_\infty \int_{-\infty}^x v \pi H dx + 2 \int_0^H u^2 \pi r dy = -\frac{D}{\rho}$$

Combinando ambas ecuaciones :

$$\left. \begin{aligned} -\bar{U}_\infty \frac{\pi(2H)^2}{4} + 2 \int_{-\infty}^x v \pi H dx + 2 \int_0^H u \pi r dr &= 0 \\ -\bar{U}_\infty \frac{\pi(2H)^2}{4} + 2\bar{U}_\infty \int_{-\infty}^x v \pi H dx + 2 \int_0^H u^2 \pi r dr &= -\frac{D}{\rho} \end{aligned} \right\} \rightarrow \int_0^H \left( u - \frac{u^2}{\bar{U}_\infty} \right) \pi r dr = \frac{D}{2\rho \bar{U}_\infty} \longrightarrow$$

$$\longrightarrow \int_0^H u \left( 1 - \frac{u}{\bar{U}_\infty} \right) \pi r dr = \frac{D}{2\rho \bar{U}_\infty} \longrightarrow \int_0^H u (\bar{U}_\infty - u) \pi r dr = \frac{D}{2\rho} \longrightarrow$$

$$\longrightarrow \boxed{\frac{D}{\rho} = 2 \int_0^H u (\bar{U}_\infty - u) \pi r dr = 2 \int_0^H u \tilde{u} \pi r dr}$$

En el exterior de la estela  $(\bar{U}_\infty - u) \rightarrow 0$ , de modo que no contribuye a la integral anterior ( $H \rightarrow \infty$ ) :

$$\forall x \in \text{estela lejana} : \int_0^\infty u \tilde{u} \pi r dr = \int_0^\infty (\bar{U}_\infty - \cancel{\tilde{u}}) \tilde{u} \pi r dr \approx \int_0^\infty \bar{U}_\infty \tilde{u} \pi r dr = \frac{D}{2\rho} \longrightarrow \boxed{\frac{D}{\rho \bar{U}_\infty} = 2 \int_0^\infty \tilde{u} \pi r dr}$$

Como consecuencia de esta disparidad de escalas podemos linearizar la Ecuación:

$$\underbrace{u \frac{\partial \tilde{u}}{\partial x}}_{\sim \bar{U}_\infty \frac{\tilde{u}_c}{x}} + \cancel{\underbrace{u \frac{\partial \tilde{u}}{\partial r}}_{\sim \tilde{u}_c \frac{\tilde{u}_c}{\delta}}} = \underbrace{\frac{1}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right)}_{\sim \sqrt{\frac{\tilde{u}_c}{\delta^2}}}$$

$$\left| \frac{1}{\tilde{u}_c} \right| \sim \frac{\sqrt{x}}{\bar{U}_\infty} \left( \frac{x}{\delta} \right)^2 \sim Re_x^{-1} \left( \frac{\delta}{x} \right)^{-2}$$

$$\left| \frac{x}{\bar{U}_\infty \tilde{u}_c} \right| \sim \frac{\tilde{u}_c^2}{x} \frac{x}{\bar{U}_\infty \tilde{u}_c} \sim \frac{\tilde{u}_c}{\bar{U}_\infty} \ll 1$$

$$\frac{\delta}{x} \sim Re_x^{-1/2} \ll 1$$

La ecuación queda :

$$u \frac{\partial \tilde{u}}{\partial x} = \frac{1}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right) \longrightarrow \bar{U}_\infty \left( 1 - \cancel{\frac{\tilde{u}}{\bar{U}_\infty}} \right) \frac{\partial \tilde{u}}{\partial x} = \frac{1}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right) \longrightarrow \boxed{\bar{U}_\infty \frac{\partial \tilde{u}}{\partial x} \approx \frac{1}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right)}$$

En cuanto a las condiciones de contorno:

- Simetría:  $\tilde{u}(r) = \tilde{u}(-r)$
- $\frac{\partial \tilde{u}}{\partial r} \Big|_{r=0} = \lim_{\Delta r \rightarrow 0} \frac{\tilde{u}(\Delta r/2) - \tilde{u}(-\Delta r/2)}{\Delta r} = 0 \longrightarrow r=0 : \frac{\partial \tilde{u}}{\partial r} = 0$

■ En  $r \rightarrow \infty$  la estrella no ejerce ninguna influencia:  $r \rightarrow \infty : u \rightarrow U_\infty \longrightarrow r \rightarrow \infty : \tilde{u} \rightarrow 0$

- Además sabemos que  $\forall x \in$  estrella lejana:  $\frac{D}{2\rho U_\infty} = \int_0^\infty \tilde{u} \pi r dr$

El problema queda:

$$U_\infty \frac{\partial \tilde{u}}{\partial x} \approx \frac{\sqrt{J}}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right)$$

$$r=0 : \frac{\partial \tilde{u}}{\partial r} = 0$$

$$r \rightarrow \infty : \tilde{u} \rightarrow 0$$

$$\forall x : \frac{D}{2\rho U_\infty} = \int_0^\infty \tilde{u} \pi r dr$$

Eliminamos  $\sqrt{J}$ ;  $U_\infty$

$$x, \sqrt{\frac{U_\infty}{J}} r, \tilde{u}$$

$$\frac{\partial \tilde{u}}{\partial x} \approx \frac{\sqrt{J}}{\sqrt{\frac{U_\infty}{J}} r} \frac{\partial}{\partial \left( \sqrt{\frac{U_\infty}{J}} r \right)} \left[ \sqrt{\frac{U_\infty}{J}} r \frac{\partial \tilde{u}}{\partial \left( \sqrt{\frac{U_\infty}{J}} r \right)} \right]$$

$$\sqrt{\frac{U_\infty}{J}} r = 0 : \frac{\partial \tilde{u}}{\partial \left( \sqrt{\frac{U_\infty}{J}} r \right)} = 0$$

$$\sqrt{\frac{U_\infty}{J}} r \rightarrow \infty : \tilde{u} \rightarrow 0$$

$$\forall x : \frac{D}{2\rho J} = I = \int_0^\infty \tilde{u} \pi \sqrt{\frac{U_\infty}{J}} r d\left(\sqrt{\frac{U_\infty}{J}} r\right)$$

Ecuaciones de dimensiones:

$$[x] = L$$

$$\left[ \sqrt{\frac{U_\infty}{J}} r \right] = \left( \frac{L T^{-1}}{L^2 T^{-1}} \right)^{1/2} L = L^{1/2}$$

$$[\tilde{u}] = L T^{-1}$$

$$[I] = \left[ \frac{D}{2\rho J} \right] = \left[ \frac{U_\infty^2 r^2}{J} \right] = L^2 T^{-1}$$

2 magnitudes  
dimensionalmente  
independientes

$$x, I$$

Adimensionalizamos :

$$\tilde{u} \longrightarrow \frac{\tilde{u}}{x^\alpha I^\beta} \longrightarrow \frac{LT^{-1}}{L^\alpha L^{2\beta} T^{-\beta}} \longrightarrow \begin{cases} \beta = 1 \\ 1 = \alpha + 2\beta \end{cases} \longrightarrow \begin{cases} \alpha = -1 \\ \beta = 1 \end{cases} \longrightarrow \frac{\tilde{u} x}{I}$$

$$\sqrt{\frac{U_{\infty}}{J}} r \longrightarrow \left[ \sqrt{\frac{U_{\infty}}{J}} r \right] = L^{1/2} \longrightarrow \text{Se adimensionaliza con } x^{1/2} \longrightarrow \sqrt{\frac{U_{\infty}}{Jx}} r = \eta$$

Entonces :

$$\boxed{\frac{\tilde{u} x}{I} = f(\eta), \text{ con } \eta = \sqrt{\frac{U_{\infty}}{Jx}} r} \quad \begin{aligned} &\rightarrow \tilde{u} = \frac{I}{x} f(\eta) \\ &\rightarrow \frac{\partial \eta}{\partial x} = -\frac{\eta}{2x} \quad \frac{\partial \eta}{\partial y} = \sqrt{\frac{U_{\infty}}{Jx}} \end{aligned}$$

Si recordamos, la ECDM quedaba reducida a :  $U_{\infty} \frac{\partial \tilde{u}}{\partial x} \approx \frac{J}{r} \frac{\partial}{\partial r} \left( \sqrt{r} \frac{\partial \tilde{u}}{\partial r} \right)$

$$\boxed{\frac{\partial \tilde{u}}{\partial x}}$$

$$\frac{\partial \tilde{u}}{\partial x} = \frac{\partial}{\partial x} \left[ \frac{I}{x} f(\eta) \right] = -\frac{I}{x^2} f - \frac{I}{x} f' \frac{\eta}{2x} \longrightarrow \boxed{\frac{\partial \tilde{u}}{\partial x} = -\frac{I}{x^2} \left( f + \frac{1}{2} \eta f' \right)}$$

$$\boxed{\frac{\partial \tilde{u}}{\partial r}}$$

$$\frac{\partial \tilde{u}}{\partial r} = \frac{\partial}{\partial r} \left[ \frac{I}{x} f(\eta) \right] = \frac{I}{x} f' \sqrt{\frac{U_{\infty}}{Jx}} = \frac{I}{x} \sqrt{\frac{U_{\infty}}{J}} f'$$

$$\boxed{\frac{\partial}{\partial r} \left( r \frac{\partial \tilde{u}}{\partial r} \right)}$$

$$\frac{\partial}{\partial r} \left( r \frac{\partial \tilde{u}}{\partial r} \right) = \frac{\partial \tilde{u}}{\partial r} + r \frac{\partial^2 \tilde{u}}{\partial r^2} = \frac{I}{x} \sqrt{\frac{U_{\infty}}{Jx}} f' + \cancel{\sqrt{\frac{Jx}{U_{\infty}}} \eta \frac{I}{x} \sqrt{\frac{U_{\infty}}{Jx}} f'' \sqrt{\frac{U_{\infty}}{Jx}}} \longrightarrow$$

$$\longrightarrow \boxed{\frac{\partial}{\partial r} \left( r \frac{\partial \tilde{u}}{\partial r} \right) = \frac{I}{x} \sqrt{\frac{U_{\infty}}{Jx}} \left( f' + \eta f'' \right)}$$

Sustituyendo:

$$-\frac{I}{x^2} \left( f + \frac{1}{2} \eta f' \right) \approx \frac{v}{r} \frac{I}{x} \sqrt{\frac{U_\infty}{J_x}} \left( f' + \eta f'' \right)$$

$$-\frac{I}{x^2} \left( f + \frac{1}{2} \eta f' \right) \approx \cancel{\sqrt{\frac{U_\infty}{J_x}}} \frac{1}{\eta} \frac{I}{x} \sqrt{\frac{U_\infty}{J_x}} \left( f' + \eta f'' \right)$$

$$-\frac{U_\infty I}{x^2} \eta \left( f + \frac{1}{2} \eta f' \right) \approx \frac{U_\infty I}{x^2} \left( f' + \eta f'' \right)$$

$$-\eta f - \frac{\eta^2}{2} f' \approx f' + \eta f''$$

$$\eta f'' + \left( 1 + \frac{\eta^2}{2} \right) f' + \eta f \approx 0$$

$$\underbrace{f'' + \frac{1}{\eta} f' + \frac{\eta}{2} f'}_{\frac{1}{\eta} (\eta f)'} + f \approx 0$$

$$\frac{1}{\eta} (\eta f)' + \frac{\eta}{2} f' + f \approx 0$$

En cuanto a las condiciones de contorno:

$$r=0 : \frac{\partial \tilde{u}}{\partial r} = 0 \longrightarrow \eta = 0 : f' = 0 \quad (1)$$

$$r \rightarrow \infty : \tilde{u} \rightarrow 0 \longrightarrow \eta \rightarrow \infty : f = 0 \quad (2)$$

Si necesitamos que  $\int_0^\infty f d\eta$  esté acotada,

se debe satisfacer  $\eta \rightarrow \infty : f \rightarrow 0$ ,

colapsando las condiciones (2) y (3)

$$\forall x : \frac{D}{2 \rho U_\infty} = \int_0^\infty \tilde{u} \pi r dr \longrightarrow I \frac{v}{U_\infty} = \int_0^\infty \frac{I}{x} f(\eta) \pi \sqrt{\frac{v x}{U_\infty}} \eta d\left(\sqrt{\frac{v x}{U_\infty}} \eta\right) \longrightarrow \int_0^\infty \eta f(\eta) d\eta = \frac{1}{\pi}$$

(3)

Por tanto, el problema queda:

$$\frac{1}{\eta} (\eta f)' + \frac{\eta}{2} f' + f \approx 0$$

$$\eta = 0 : f' = 0$$

$$\forall x : \int_0^\infty \eta f(\eta) d\eta = \frac{1}{\pi}$$

DESHACE LA HOMOGENEIDAD

"D" RESPONSABLE DE LA APARICIÓN DE LA ESTELA

Integración:

$$\frac{1}{t} \left( t f' + \frac{t^2}{2} f \right)' \approx 0$$



$$t f' + \frac{t^2}{2} f \approx C_1 \quad \xrightarrow{t=0 : f=0} C_1 \approx 0$$

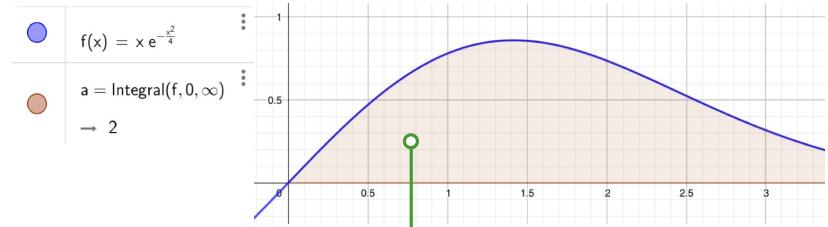


$$t f' + \frac{t^2}{2} f \approx 0$$

$$f' + \frac{t}{2} f \approx 0$$



$$\int \frac{f'}{f} df \approx -\frac{1}{2} \int t dt$$

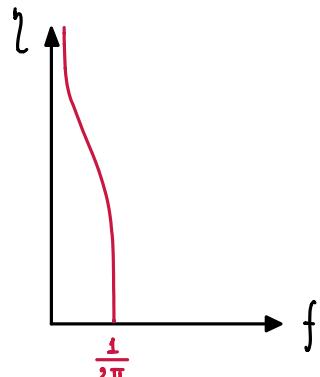


$$\ln f + k_1 \approx -\frac{t^2}{4} + k_2$$

$$f \approx C_2 \exp\left(-\frac{t^2}{4}\right) \quad \xrightarrow{\int_0^\infty t f dt = \frac{1}{\pi}} \quad C_2 \approx \frac{\frac{1}{\pi}}{\boxed{\int_0^\infty t \exp\left(-\frac{t^2}{4}\right) dt}} = \frac{1}{2\pi}$$



$$f(t) \approx \frac{1}{2\pi} \exp\left(-\frac{t^2}{4}\right)$$



En cuanto a "u":

$$u = U_{\infty} - \tilde{u} = U_{\infty} - \frac{\frac{D}{2\rho}x}{2\pi} \frac{1}{r^2} \exp\left(-\frac{U_{\infty}}{4\rho} r^2\right)$$

$$u = U_{\infty} - \frac{\frac{D}{\rho}}{4\pi x} \exp\left(-\frac{U_{\infty}}{4\rho} x^2\right)$$

Nota:

$$\tilde{u} \Big|_{r=0} \sim x^{-\frac{1}{2}}$$

El perfil de velocidades se va uniformizando al alejarnos del obstáculo, pero más rápidamente que en estela plana ( $\sim x^{-1/2}$ ). Es decir, el efecto del obstáculo abarca una menor distancia en 3D.